

# **A Long-Duration Balloon Experiment**

Payload for Ultrahigh Energy Observations (PUEO)

Last edited: 2025-09-24

# Overview of the Experiment

---

- PUEO is the successor of the Antarctic Impulsive Transient Antenna (ANITA) experiments (A1 to 4).

# Payload for Ultrahigh Energy Observations (PUEO) Overview of the Experiment



- PUEO is the successor of the Antarctic Impulsive Transient Antenna (ANITA) experiments (A1 to 4).
- Mainly, we aim to detect impulsive radio signals using antennas.

# Payload for Ultrahigh Energy Observations (PUEO) Overview of the Experiment



- PUEO is the successor of the Antarctic Impulsive Transient Antenna (ANITA) experiments (A1 to 4).
- Mainly, we aim to detect impulsive radio signals using antennas.
- The PUEO payload you see on the left has 96 antennas.

# Payload for Ultrahigh Energy Observations (PUEO) Overview of the Experiment

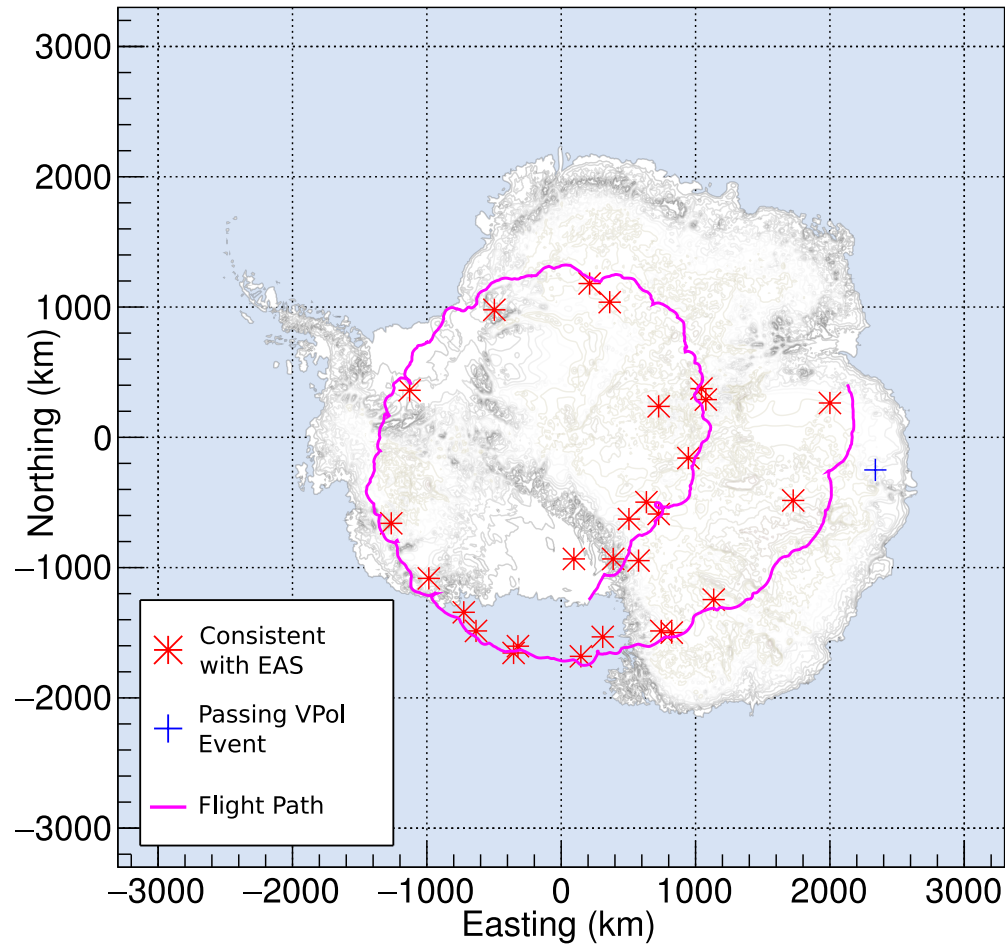


- PUEO is the successor of the Antarctic Impulsive Transient Antenna (ANITA) experiments (A1 to 4).
- Mainly, we aim to detect impulsive radio signals using antennas.
- The PUEO payload you see on the left has 96 antennas. It is to be carried by a big big balloon.



- The impulsive radio signals mentioned earlier could be produced by ultrahigh energy neutrinos (above  $10^{17}$  eV)

# The Mission

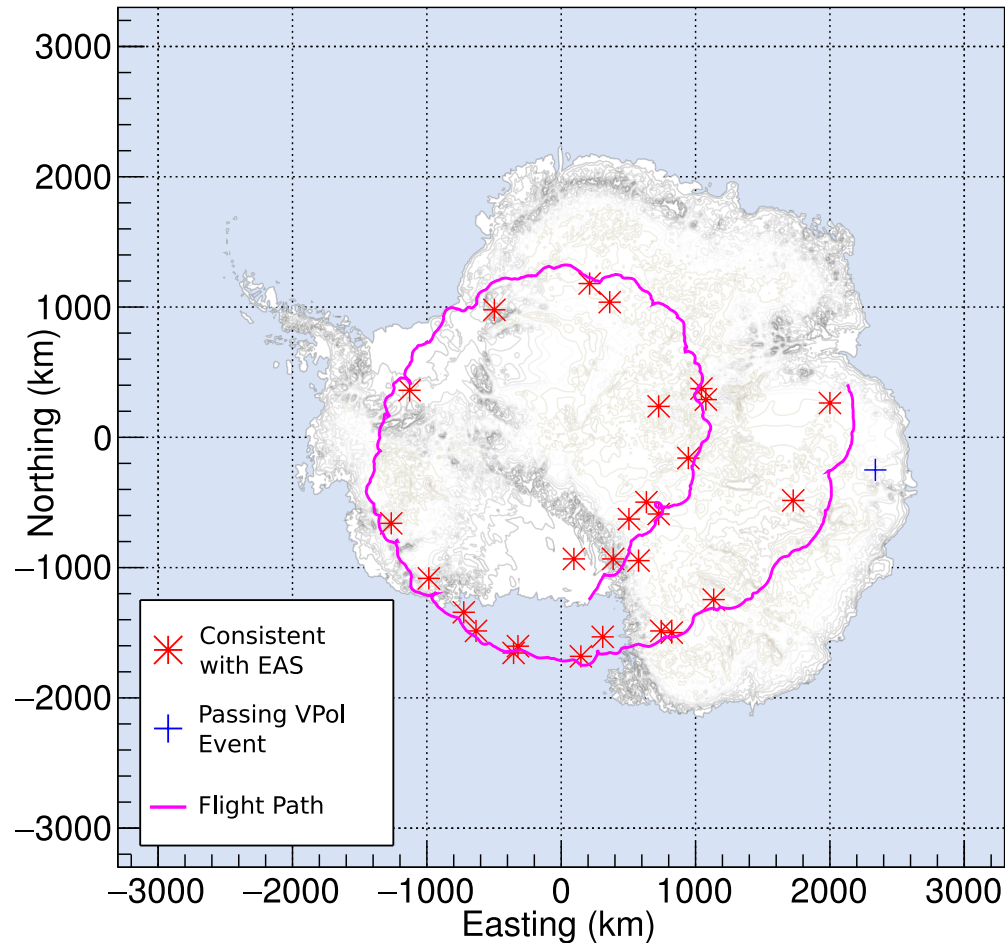


## Overview of the Experiment

- The impulsive radio signals mentioned earlier could be produced by ultrahigh energy neutrinos (above  $10^{17}$  eV)
- The balloon, carried by the summer polar vortex, circles Antarctica to look for the signals.



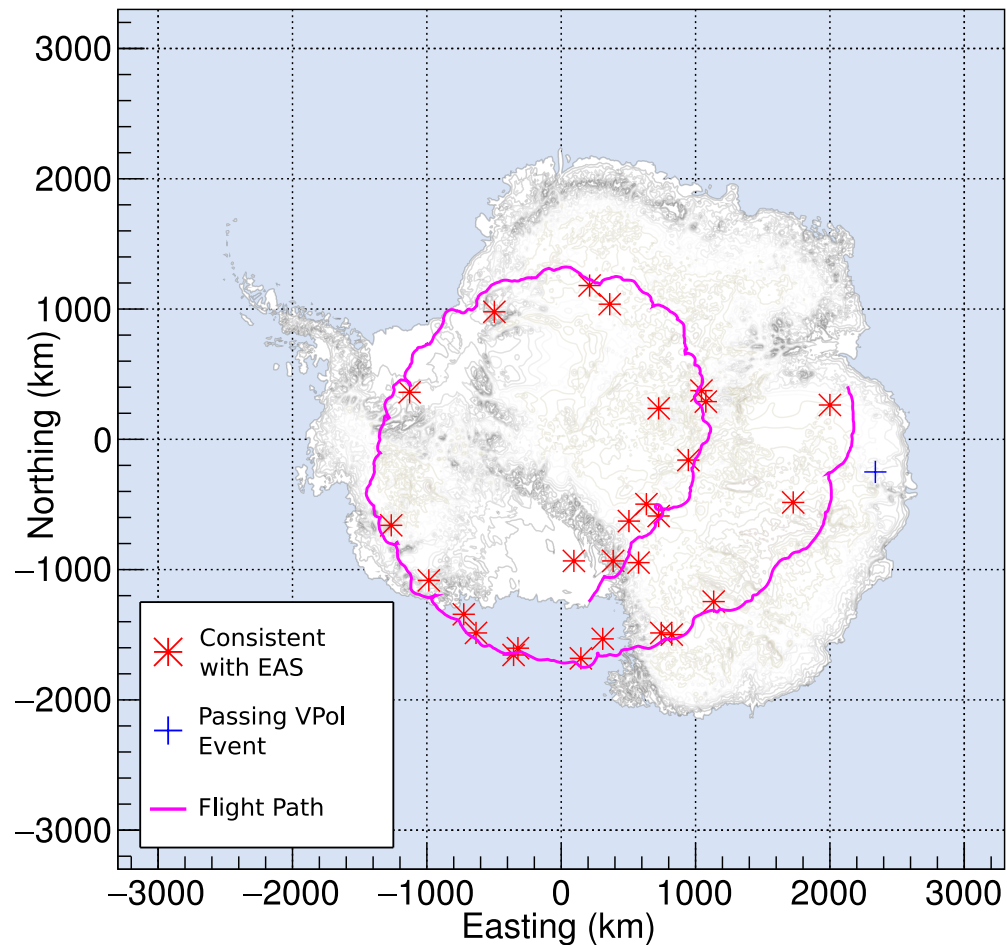
# The Mission



## Overview of the Experiment

- The impulsive radio signals mentioned earlier could be produced by ultrahigh energy neutrinos (above  $10^{17}$  eV)
- The balloon, carried by the summer polar vortex, circles Antarctica to look for the signals.
- These events ( $O[10^{-2}/\text{km}^2/\text{year}]$ ) are very rare, so it helps that the balloon flies high up there ( $\sim 35$  to  $40$  km above ice, which itself is about  $2\text{km}$  deep) and sees far away.

# The Mission



## Overview of the Experiment

- The impulsive radio signals mentioned earlier could be produced by ultrahigh energy neutrinos (above  $10^{17}$  eV)
- The balloon, carried by the summer polar vortex, circles Antarctica to look for the signals.
- These events ( $O[10^{-2}/\text{km}^2/\text{year}]$ ) are very rare, so it helps that the balloon flies high up there ( $\sim 35$  to  $40$  km above ice, which itself is about  $2\text{ km}$  deep) and sees far away.
- A3 at any one time surveyed a  $\sim 600\text{-km}$ -radius disk. Assuming ice depth  $\sim 1$  km, the detection volume is about  $1$  million  $\text{km}^3$ .

# Why Neutrino?

---

- $\nu$  does not interact electromagnetically. This makes detection challenging, but it also makes  $\nu$  very useful for probing distant objects, as they travel in a straight line in electromagnetic fields, pointing back to their sources.

- $\nu$  does not interact electromagnetically. This makes detection challenging, but it also makes  $\nu$  very useful for probing distant objects, as they travel in a straight line in electromagnetic fields, pointing back to their sources.
- On the other hand, there exists a horizon for ultrahigh energy photons. For example,  $\gamma_{\text{UHE}} \gamma_{\text{CMB}}$  interact with each other and produces lower energy photons.
- In other words, we can “see” the world through neutrinos, provided that we can (1) detect them, and (2) figure out their directions.

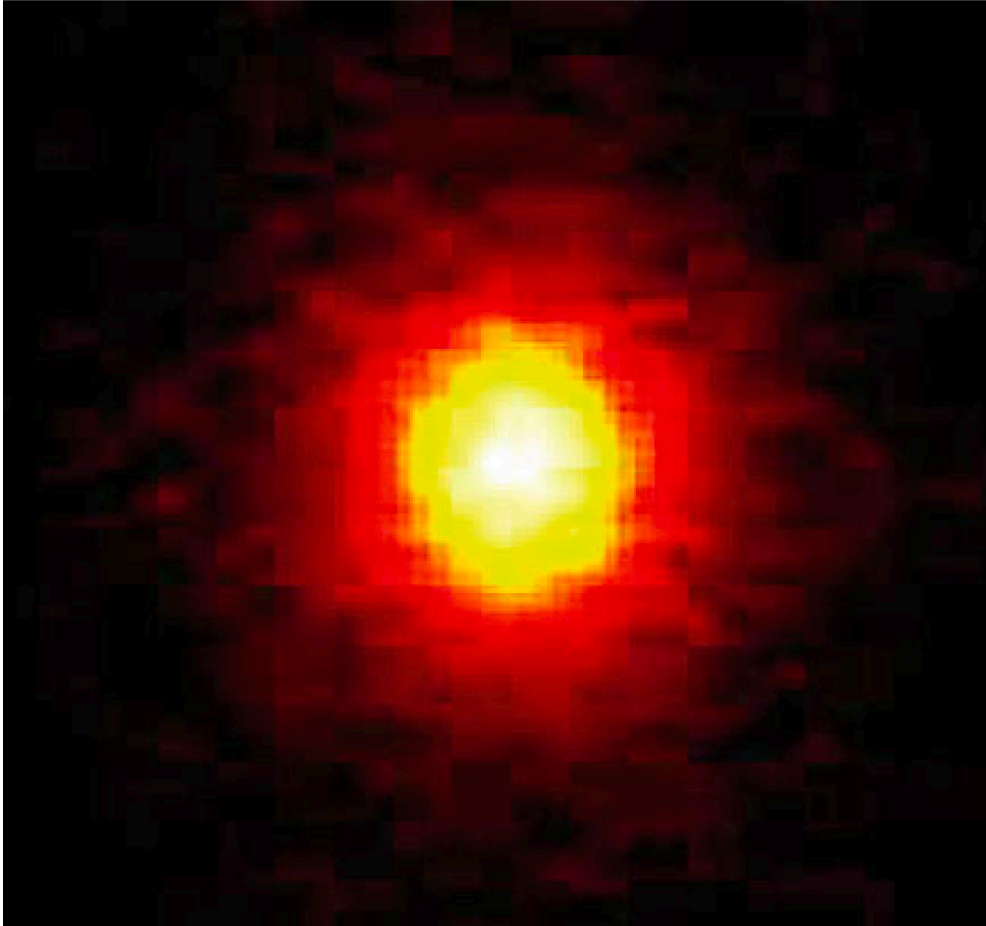


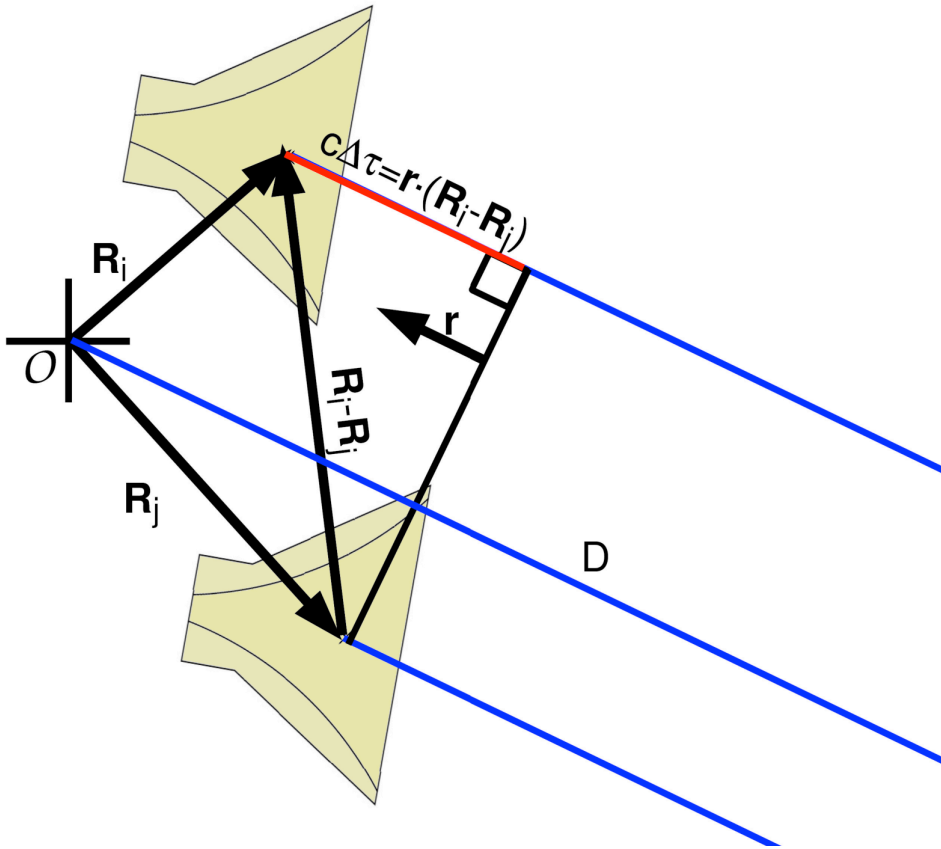
Figure 1: A picture of our sun, taken by the Super-Kamiokande neutrino detector

- $\nu$  does not interact electromagnetically. This makes detection challenging, but it also makes  $\nu$  very useful for probing distant objects, as they travel in a straight line in electromagnetic fields, pointing back to their sources.
- On the other hand, there exists a horizon for ultrahigh energy photons. For example,  $\gamma_{\text{UHE}} \gamma_{\text{CMB}}$  interact with each other and produces lower energy photons.
- In other words, we can “see” the world through neutrinos, provided that we can (1) detect them, and (2) figure out their directions.

# Direction Reconstruction: PUEO

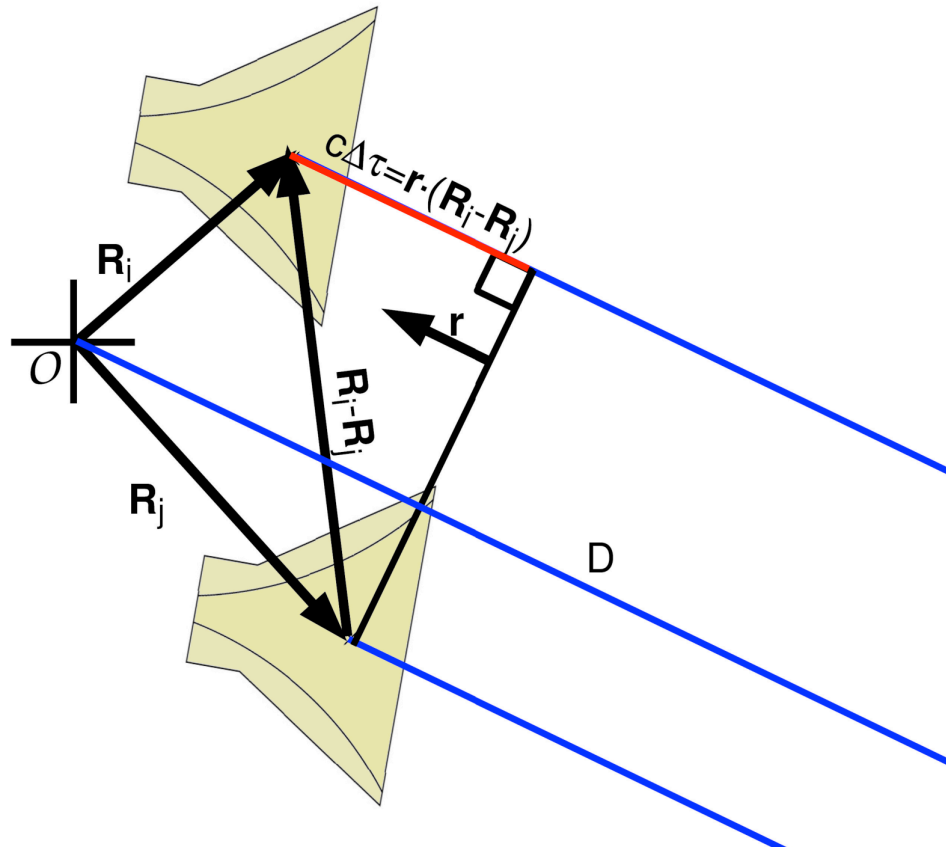
---

- Consider two antennas
  - $\mathbf{R}_i$  and  $\mathbf{R}_j$  denote the locations of a generic pair of antennas.
  - $\mathbf{R}_i - \mathbf{R}_j$  denotes the displacement vector.



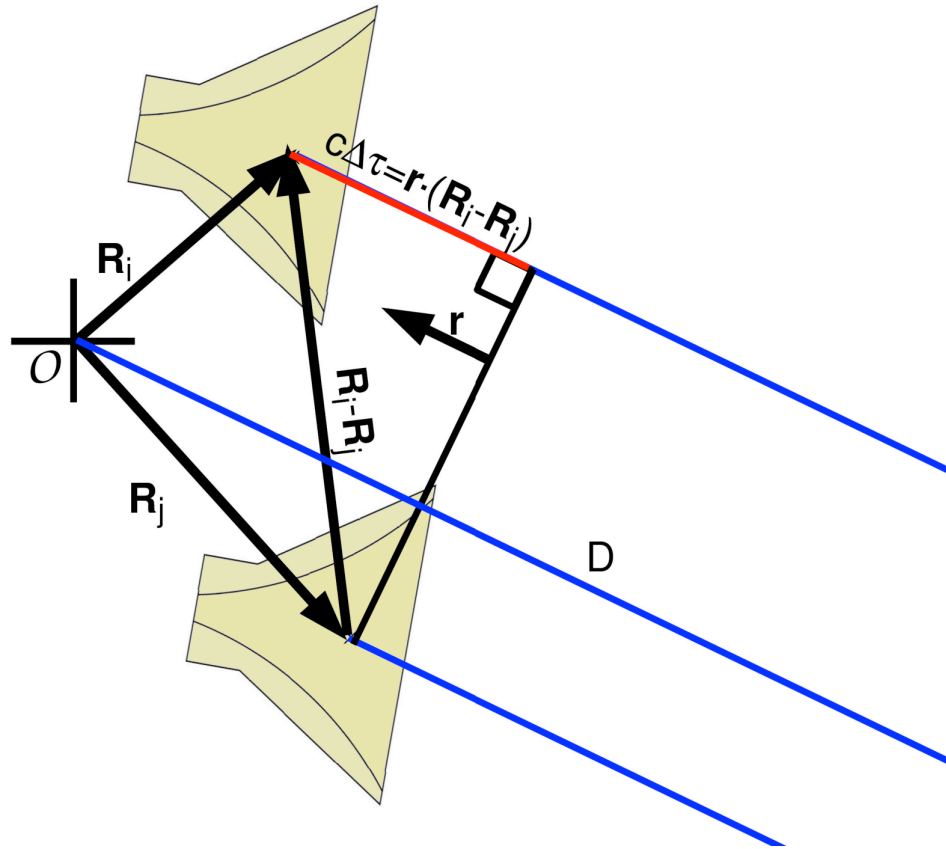


# Signal Delay: Some Trig



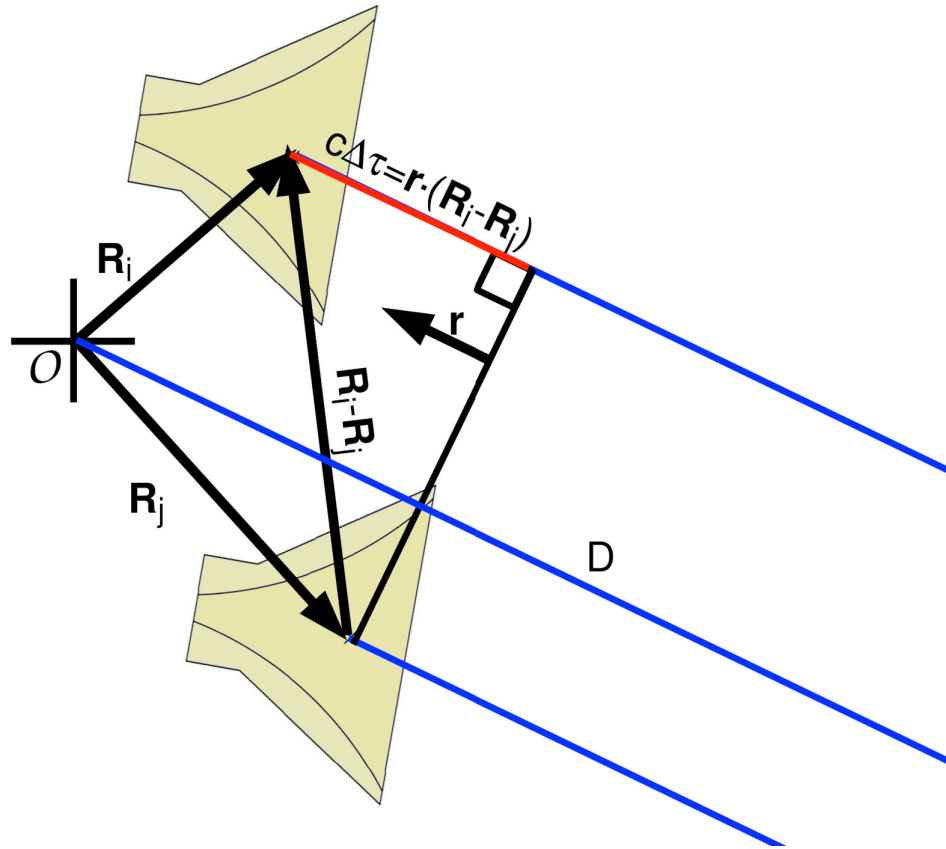
- Consider two antennas
  - $\mathbf{R}_i$  and  $\mathbf{R}_j$  denote the locations of a generic pair of antennas.
  - $\mathbf{R}_i - \mathbf{R}_j$  denotes the displacement vector.
- $\hat{\mathbf{r}}$ : signal direction (assume plane wave)

# Signal Delay: Some Trig

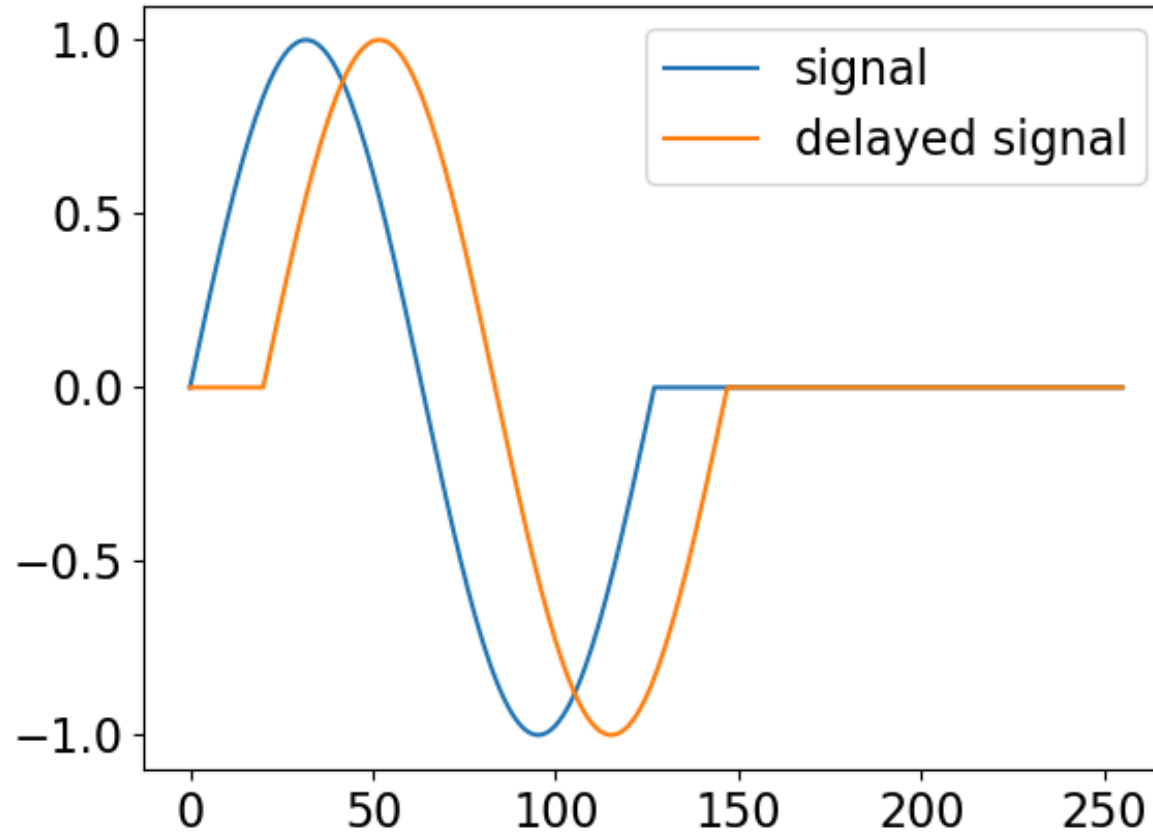


- Consider two antennas
  - $R_i$  and  $R_j$  denote the locations of a generic pair of antennas.
  - $R_i - R_j$  denotes the displacement vector.
- $\hat{r}$ : signal direction (assume plane wave)
- $c \cdot \Delta\tau$ : the extra distance that the signal needs to travel to hit the second antenna

# Signal Delay: Some Trig

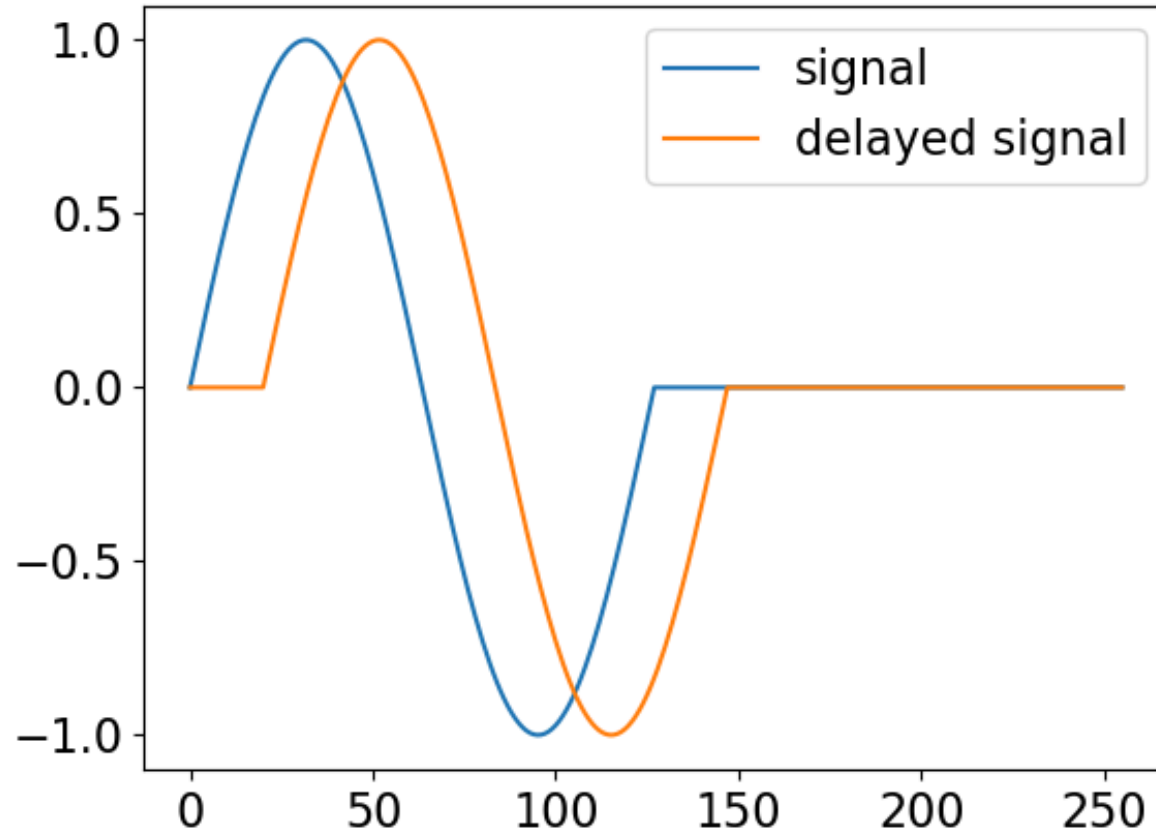


- Consider two antennas
  - $R_i$  and  $R_j$  denote the locations of a generic pair of antennas.
  - $R_i - R_j$  denotes the displacement vector.
- $\hat{r}$ : signal direction (assume plane wave)
- $c \cdot \Delta\tau$ : the extra distance that the signal needs to travel to hit the second antenna
- Evidently we can compute the extra distance by taking a dot product



- Now consider what shows up on our computer

# Signal Delay: Cross-Correlation

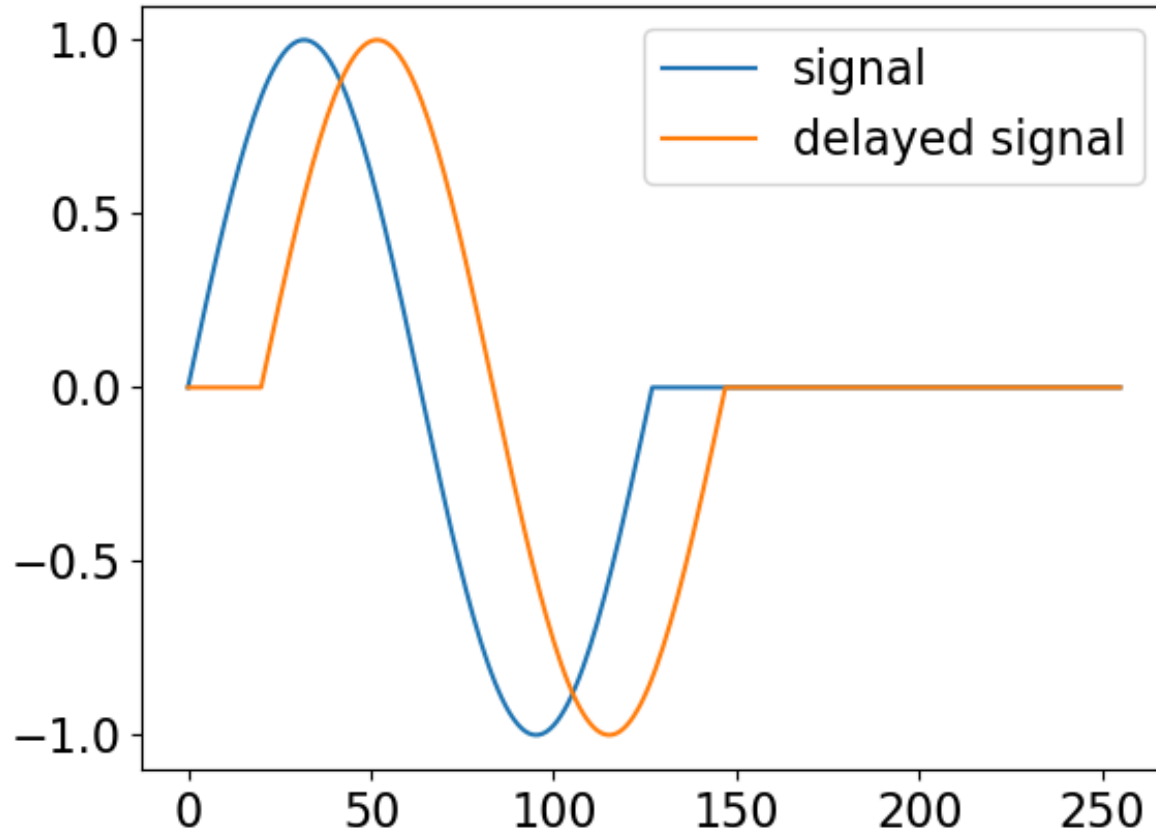


## Direction Reconstruction: PUEO

- To measure the time delay, we compute the **cross-correlation**.

- Now consider what shows up on our computer

# Signal Delay: Cross-Correlation

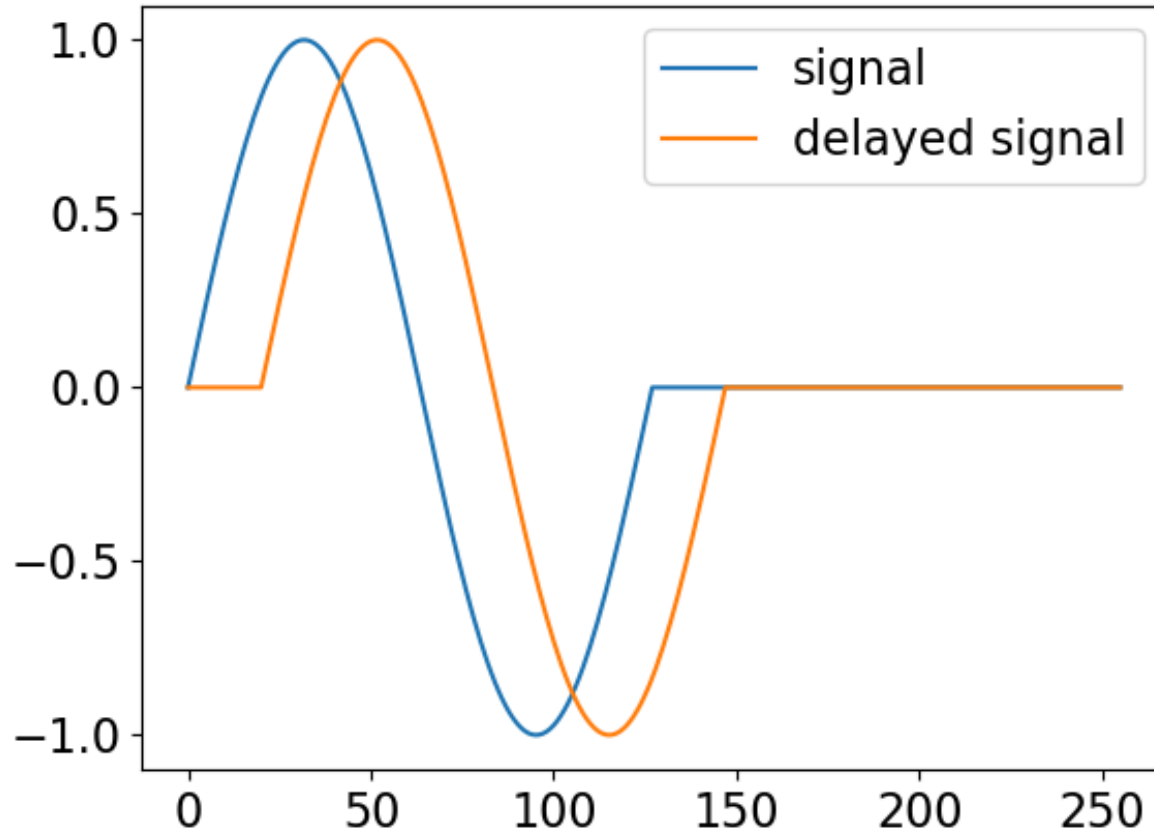


## Direction Reconstruction: PUEO

- To measure the time delay, we compute the **cross-correlation**.
- Step 1: sum the arrays `signal_1[1024]` and `signal_2[1024]` together.

- Now consider what shows up on our computer

# Signal Delay: Cross-Correlation

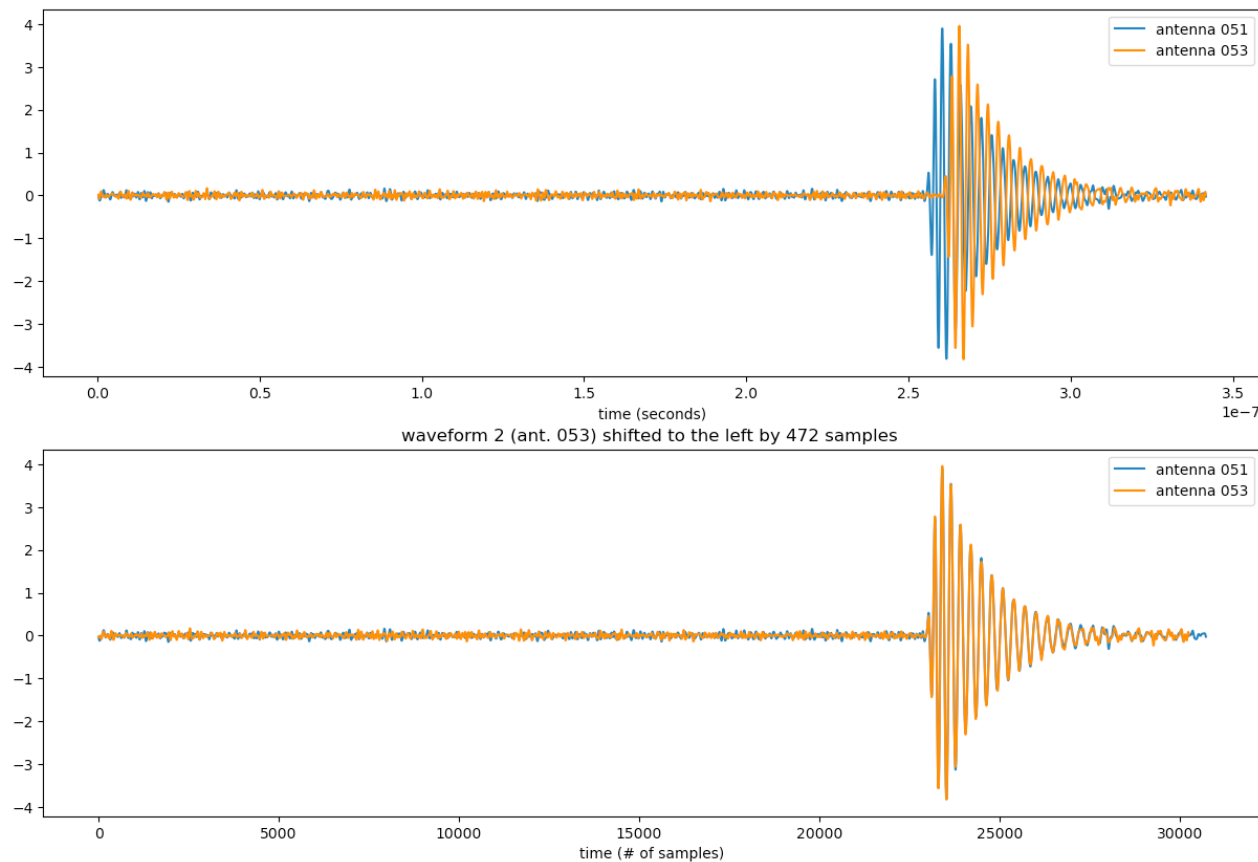


- Now consider what shows up on our computer

## Direction Reconstruction: PUEO

- To measure the time delay, we compute the **cross-correlation**.
- Step 1: sum the arrays `signal_1[1024]` and `signal_2[1024]` together.
- Step 2: shift `signal_2[1024]` to the left (or right) a little bit, and sum again.

# Signal Delay: Cross-Correlation

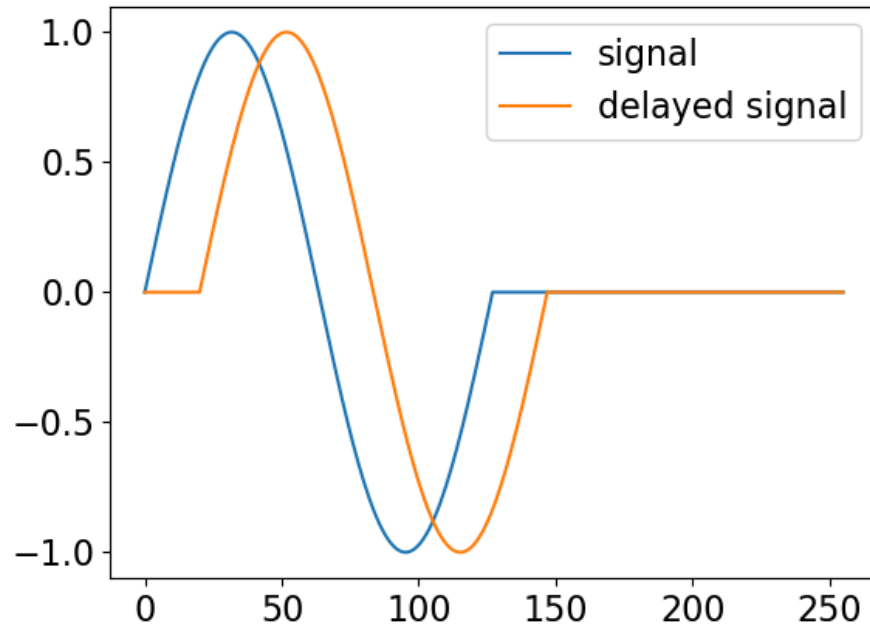


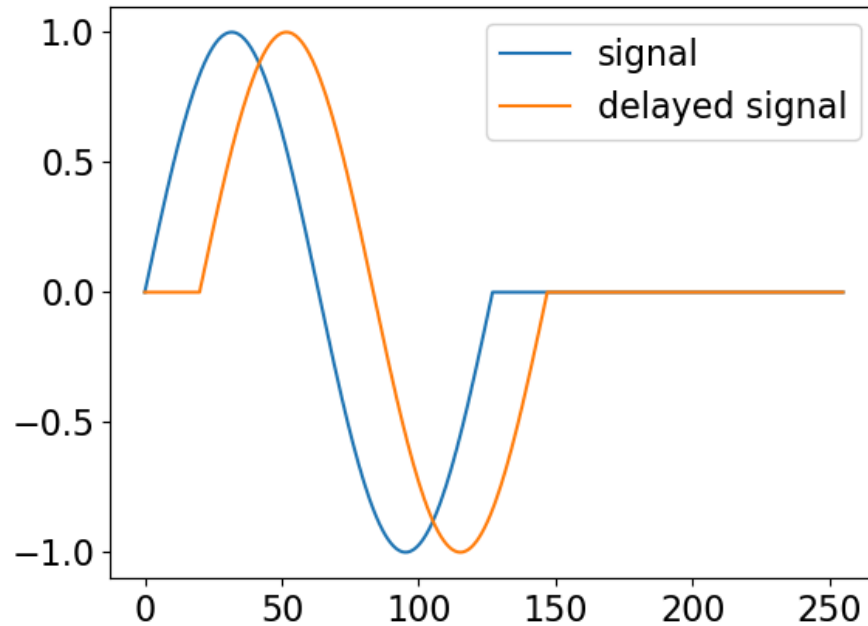
- Now consider what shows up on our computer

## Direction Reconstruction: PUEO

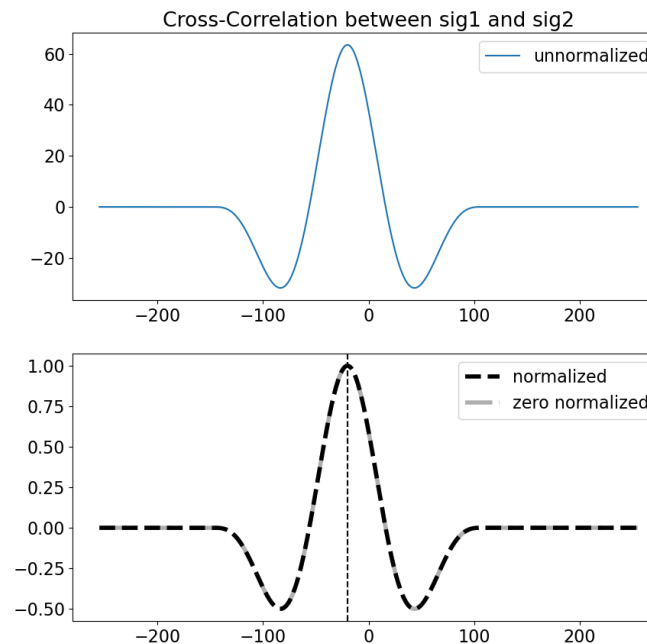
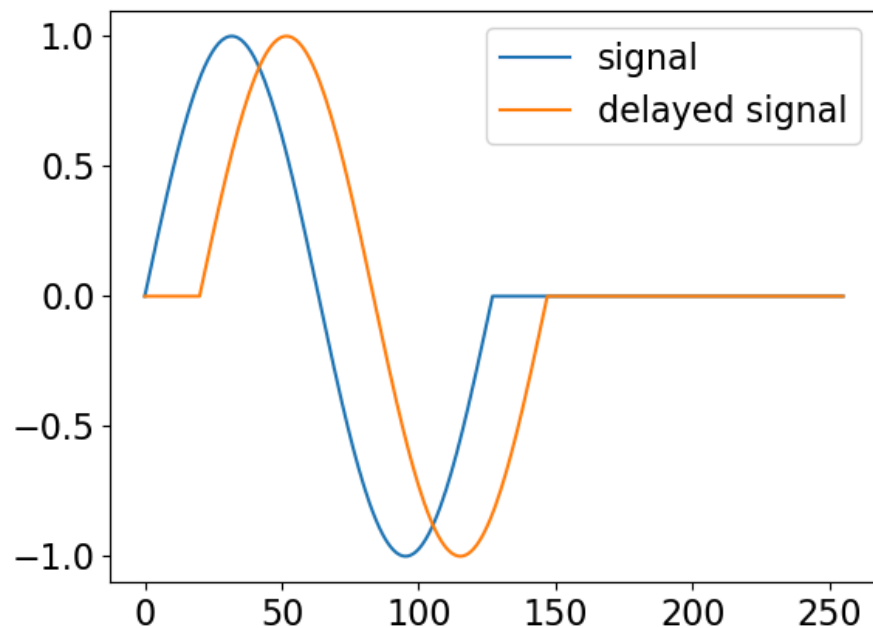
- To measure the time delay, we compute the **cross-correlation**.
- Step 1: sum the arrays `signal_1[1024]` and `signal_2[1024]` together.
- Step 2: shift `signal_2[1024]` to the left (or right) a little bit, and sum again.
- Eventually, when the two signals are on top of each other, the sum would be the largest.



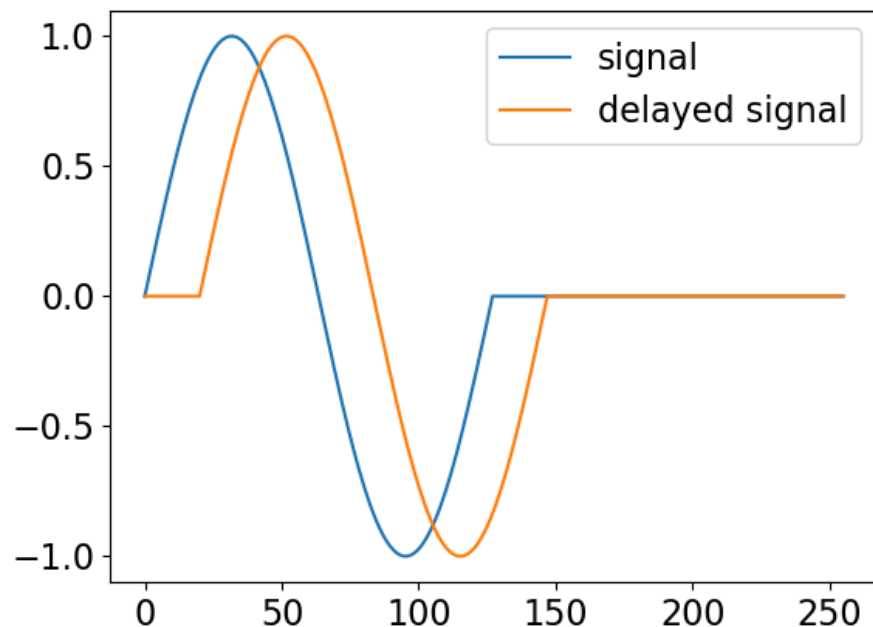




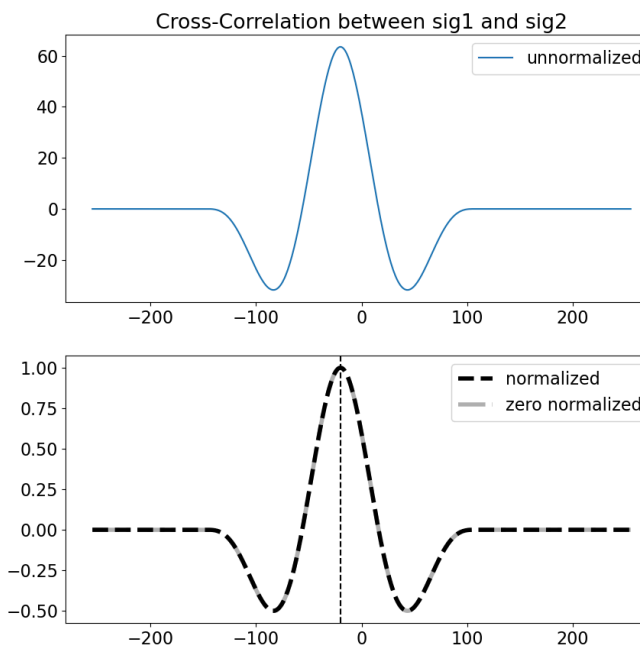
If we cross correlate the two signals above, we would see that the correlation score peaks at **-20**.



If we cross correlate the two signals above, we would see that the correlation score peaks at **-20**.



If we cross correlate the two signals above, we would see that the correlation score peaks at **-20**.



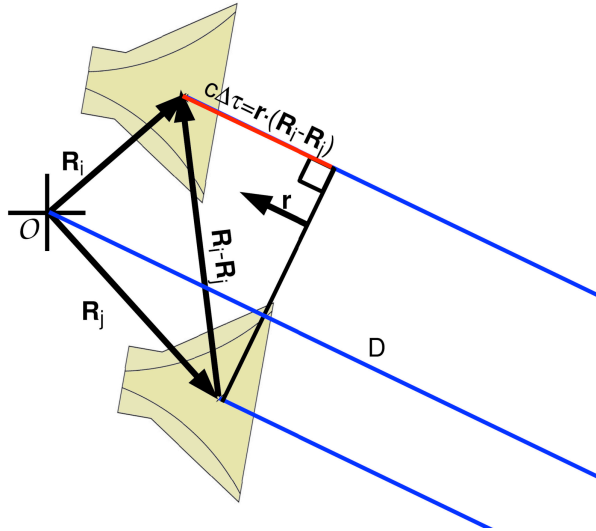
This means if we shift the **delayed signal** to the left by 20, the two signals would align.

- Take away: by (blindly) shifting and summing, we will eventually get a peak in correlation scores
- Peak correlation means we have found the correct time delay.

For some signal, we blindly “guess” where it could come from; call this direction  $\hat{\mathbf{r}}_{\text{guess}}$

# Correlation Map

For some signal, we blindly “guess” where it could come from; call this direction  $\hat{\mathbf{r}}_{\text{guess}}$



Then we take the dot product to “figure out” the time delay corresponding to that guess.

For some signal, we blindly “guess” where it could come from; call this direction  $\hat{\mathbf{r}}_{\text{guess}}$

Then we take the dot product to “figure out” the time delay corresponding to that guess.

Should this guess  $\hat{\mathbf{r}}_{\text{guess}}$  be correct, the time delay we compute would be correct, so we would be shifting `signal_2[1024]` by the correct amount such that we get a peak correlation.

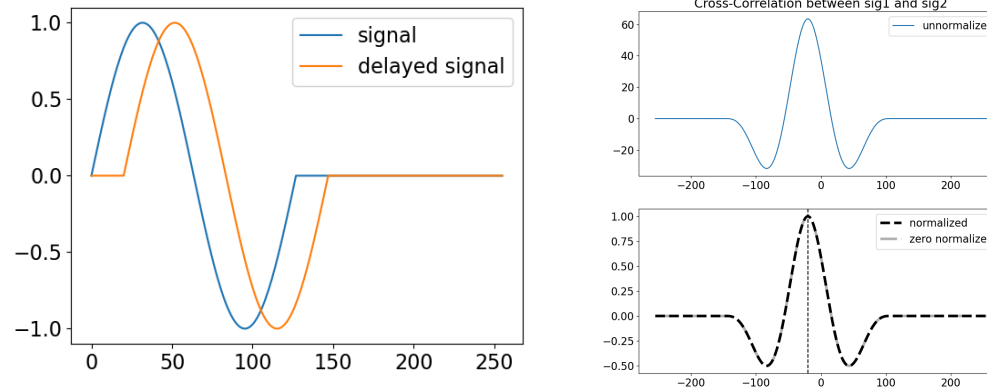


# Correlation Map

For some signal, we blindly “guess” where it could come from; call this direction  $\hat{\mathbf{r}}_{\text{guess}}$

Then we take the dot product to “figure out” the time delay corresponding to that guess.

Should this guess  $\hat{\mathbf{r}}_{\text{guess}}$  be correct, the time delay we compute would be correct, so we would be shifting `signal_2[1024]` by the correct amount such that we get a peak correlation.

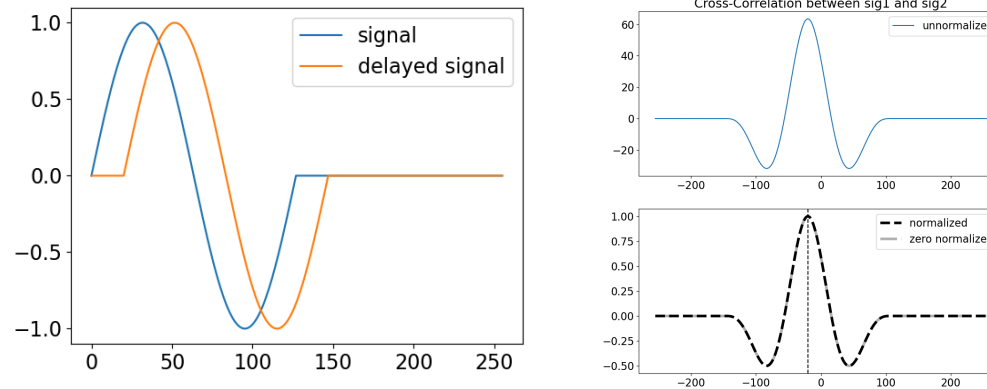


# Correlation Map

For some signal, we blindly “guess” where it could come from; call this direction  $\hat{\mathbf{r}}_{\text{guess}}$

Then we take the dot product to “figure out” the time delay corresponding to that guess.

Should this guess  $\hat{\mathbf{r}}_{\text{guess}}$  be correct, the time delay we compute would be correct, so we would be shifting `signal_2[1024]` by the correct amount such that we get a peak correlation.



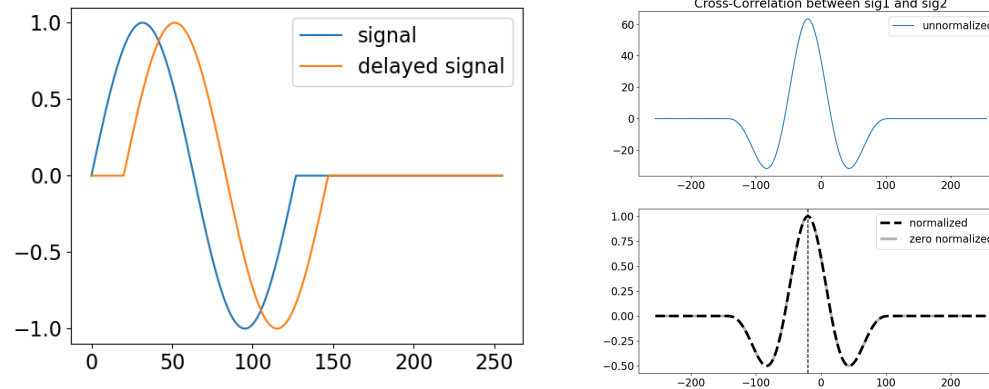
Easy as 1-2-3!

# Correlation Map

For some signal, we blindly “guess” where it could come from; call this direction  $\hat{\mathbf{r}}_{\text{guess}}$

Then we take the dot product to “figure out” the time delay corresponding to that guess.

Should this guess  $\hat{\mathbf{r}}_{\text{guess}}$  be correct, the time delay we compute would be correct, so we would be shifting `signal_2[1024]` by the correct amount such that we get a peak correlation.

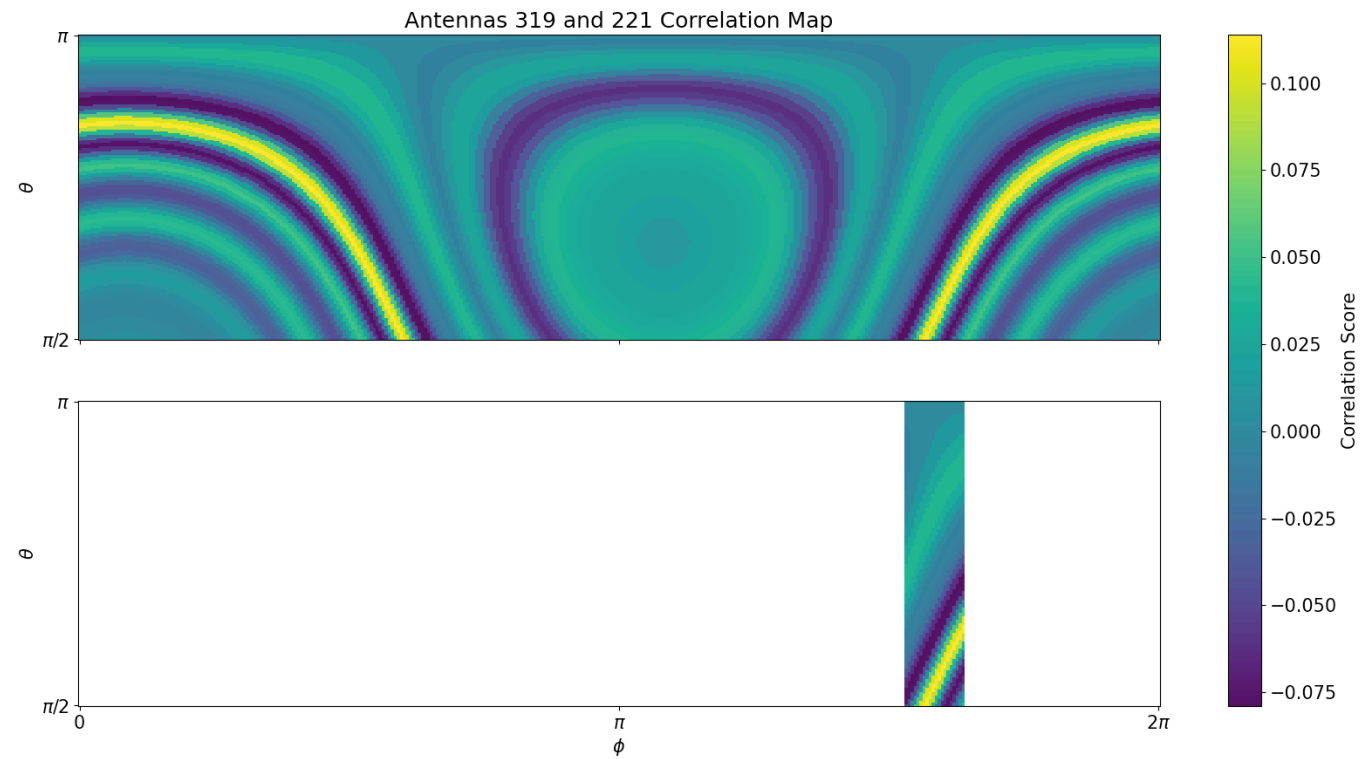


Easy as 1-2-3! (except we do this for  $O[1000]$  pairs of antennas, for each pair there are typically  $O[10000]$   $\hat{\mathbf{r}}_{\text{guess}}$ , and oh the signal length is not really 1024 because we “upsample”, but those are details.)

Anyway, the result is what we call a **correlation map** (for a single pair of antennas)

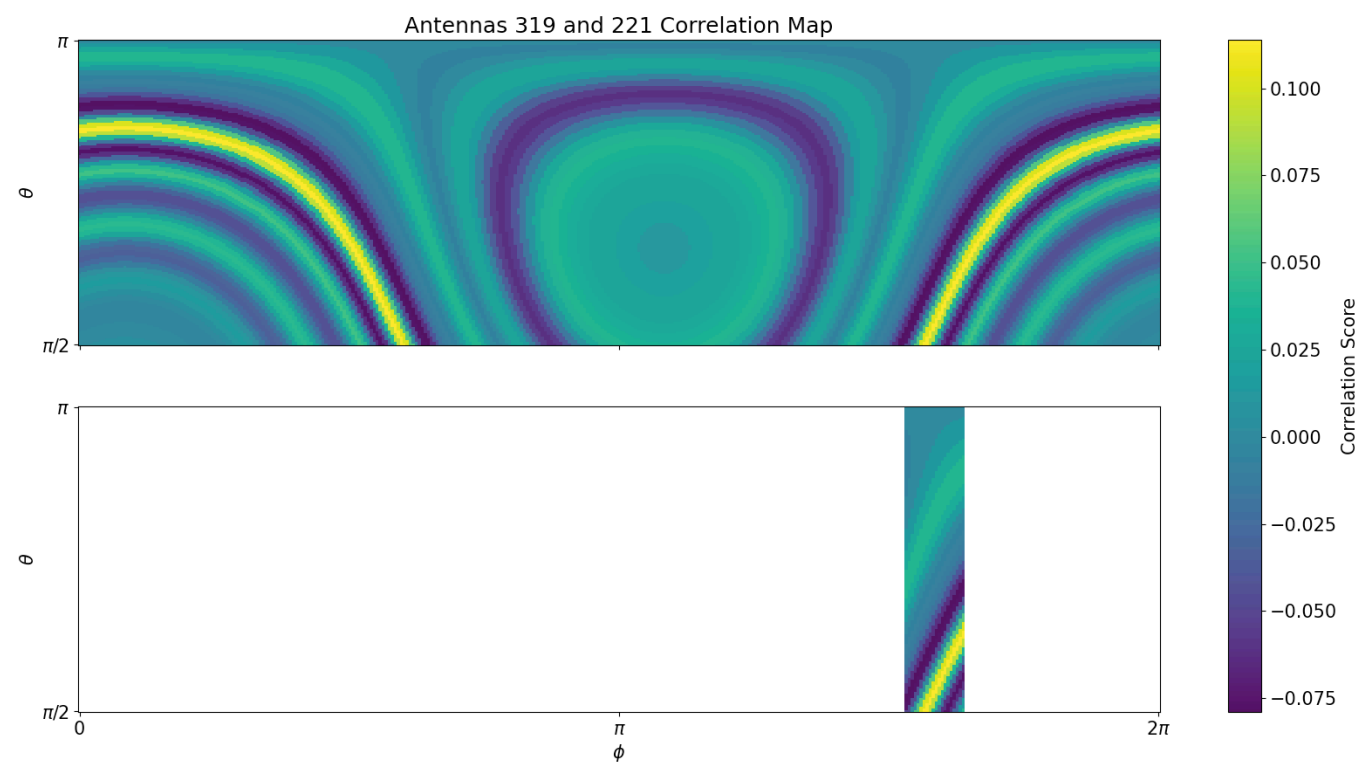
# Correlation Map

The  $\hat{\mathbf{r}}_{\text{guess}}$  mentioned earlier are the tiny direction **bins** in the map below.



# Correlation Map

You can see that for some bins the correlation is strong, meaning that the signal *could* be coming from these directions, as far as **this antenna pair** is concerned.



# Correlation Map

If we now overlay all pair-wise maps (all  $O[1000]$  antenna pairs) on top of each other:

